

# The role of geomagnetic field intensity in late Quaternary evolution of humans and large mammals

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**Key Points:** (1) The strength of the geomagnetic field is a proxy for the flux of ultra-violet radiation (UVR). (2) The disappearances of the Neanderthals and many large mammals during the Late Quaternary occur during minima in geomagnetic field strength. (3) Human phylogeny from mitochondrial DNA and Y-chromosomes can also be linked to minima in field strength, hence UVR flux.

## Abstract

It has long been speculated that biological evolution was influenced by ultra-violet radiation (UVR) reaching the Earth's surface, despite imprecise knowledge of the timing of both UVR flux and evolutionary events. The past strength of Earth's dipole field provides a proxy for UVR flux because of its role in maintaining stratospheric ozone. The timing of Quaternary evolutionary events has become better constrained by fossil finds, improved radiometric dating, use of dung fungi as proxies for herbivore populations, and improved ages for nodes in human phylogeny from human mitochondrial DNA (mtDNA) and Y-chromosomes. The demise of Neanderthals at ~41 ka can now be closely tied to the intensity minimum associated with the Laschamp magnetic excursion, and the survival of anatomically modern humans (AMHs) can be attributed to differences in the aryl hydrocarbon receptor (AhR) that has a key role in the evolutionary response to UVR flux. Fossil occurrences and dung-fungal proxies in Australia indicate that episodes of Late Quaternary extinction (LQE) of mammalian megafauna occurred close to the Laschamp and Blake magnetic excursions. Fossil and dung fungal evidence for the age of the LQE in North America (and Europe) coincide with a prominent decline in geomagnetic field intensity at ~13 ka. Over the last ~200 kyr, phylogeny based on mtDNA and Y-chromosomes in modern humans yield nodes and bifurcations in evolution corresponding to geomagnetic intensity minima which supports the proposition that UVR reaching Earth's surface influenced mammalian evolution with the loci of extinction controlled by the geometry of stratospheric ozone depletion.

## Plain Language Summary

The strength of Earth's magnetic field in the past, recorded by rocks and sediments, provides a proxy for past flux of ultra-violet radiation (UVR) to Earth's surface due to the role of the field in modulating stratigraphic ozone. About 40 thousand years ago, mammalian fossils in Australia and Eurasia record an important die-off of large mammals that included Neanderthals in Europe. In the Americas and Europe, a large

50 mammal die-off appears to have occurred ~13 thousand years ago. Both die-offs can be  
51 linked to minima in Earth's magnetic field strength implying that UVR flux variations to  
52 Earth's surface influenced mammalian evolution. For the last ~200 thousand years,  
53 estimates of the timing of branching episodes in the human evolutionary tree, from  
54 modern and fossil DNA and Y-chromosomes, can be linked to minima in field strength  
55 which implies a long-term role for UVR in human evolution. New fossil finds, improved  
56 fossil dating, knowledge of the past strength of Earth's magnetic field, and refinements  
57 in the human evolutionary tree, are sharpening the focus on a possible link between  
58 UVR arriving at the Earth's surface, magnetic field strength, and events in mammalian  
59 evolution.

60

61 **Index Terms: Paleointensity (1521), Evolutionary Geobiology (0444),**  
62 **Macropaleontology (0459)**

63

64

## 65 **1. Introduction**

66

67 The apparent spacing of mass extinction events at long (~26 Ma\*) repeat times (e.g.,  
68 Raup & Sepkoski, 1986), and the supposed role of geomagnetic polarity reversal in  
69 extinction (e.g., Raup, 1985), have received intermittent attention over the last 50 years,  
70 since early studies of Quaternary radiolarian evolution and polarity reversal in deep-sea  
71 sediments (Hays, 1971). These efforts have not resulted in significant traction for claims  
72 of a linkage between polarity reversal (with its concomitant low field intensity) and  
73 extinction or speciation, perhaps because of uncertainties in the polarity timescale  
74 itself, and in the chronology of extinction/speciation outside of the few well-  
75 documented mass extinctions. On the other hand, we now know from Quaternary  
76 studies that although polarity reversals coincided with relative paleointensity (RPI)  
77 minima, intervals between polarity reversals are also characterized by numerous RPI  
78 minima, some of which coincide with magnetic excursions (see Laj & Channell, 2015).  
79 The chronology of both the RPI record and the paleontological record remains poorly  
80 constrained, even for the Quaternary, such that a linkage between extinction and RPI  
81 minima cannot be ruled out.

82

83 The geomagnetic field helps to preserve stratospheric ozone, as well as atmospheric  
84 composition, density and oxygen levels that are vital to Earth's biosphere (Wei et al.,  
85 2014). The field shields Earth from galactic cosmic rays (GCR) and solar wind, and from  
86 harmful ultraviolet radiation (UVR) that affect the function of living systems (Belisheva  
87 et al., 2012; Mendoza & de la Pena, 2010). The demise of the Martian magnetic field,  
88 several billion years ago, is widely believed to have been the root cause for the near  
89 disappearance of the Martian atmosphere and the resulting dramatic change in the  
90 Martian environment from one featuring surface water and aqueous sedimentation to  
91 its present relative inactivity and sterility. The explosion of life in the Early Cambrian  
92 period at ~530 Ma has been associated with growth of Earth's inner core, the supposed  
93 strengthening of the dipole geomagnetic field, and the resulting thickening of Earth's

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\* The reader is referred to Aubry et al. (2009) for abbreviations denoting geological time in the past (Ma for millions of years ago, and ka for thousands of years ago) and equivalent durations (Myr and kyr, respectively).

94 atmosphere (Doglioni et al., 2016), although there is little evidence for strengthening of  
95 the geomagnetic field at this time (e.g., Biggin et al., 2015). On the other hand, the Late  
96 Ediacaran and Early Cambrian periods (~550 and ~530 Ma, respectively) may have  
97 been times of unusually high polarity reversal frequency (Pavlov & Gallet, 2001;  
98 Bazhenov et al., 2016), although precise estimates of reversal frequency are elusive due  
99 to poorly constrained age control in stratigraphic sections where the reversals were  
100 recorded. Meert et al. (2016) proposed that high reversal frequency (up to ~20  
101 reversals/Myr) at this time would have been associated with low geomagnetic field  
102 intensity that, therefore, lowered shielding from UVR, which created an evolutionary  
103 advantage for burrowing and shelled organisms. These proposals for the role of the  
104 geomagnetic field in evolution are controversial partly because of poor knowledge of  
105 the state of the geomagnetic field 500-550 million years ago. Oxygenation of the oceans  
106 and atmosphere after the Gaskiers glaciation at ~580 Ma (Canfield et al., 2007) may  
107 have been the principal driver of the Early Cambrian explosion of life, both through  
108 oxygen levels at Earth's surface and increased UVR shielding through enhanced  
109 stratospheric ozone concentrations.

110  
111 Several strategies in modern organisms reflect the evolutionary impact of UVR.  
112 Behavioral adaptations to UVR include vertical water-column migration in aquatic  
113 organisms, the presence of UVR-screening pigmentation decreasing with water depth,  
114 and complete disappearance of pigments for deep-water and cave-dwelling animals  
115 (e.g., Hessen, 2008). The red coloration of alpine plankton and the "red sweat" of the  
116 hippopotamus (Saikawa et al., 2014) are examples of evolutionary adaptation to high  
117 UVR at altitude and at low latitudes, respectively. UVR causes two classes of DNA  
118 lesions: cyclobutane pyrimidine dimers (CPDs) and 6-4 photoproducts (6-4 PPs). Both  
119 lesions distort DNA structure, introducing bends or kinks and thereby impeding  
120 transcription and replication (e.g., Clancy, 2008; Branze & Foiani, 2008). Relatively  
121 flexible areas of the DNA double helix are most susceptible to damage. One "hot spot"  
122 for UV-induced damage is found within a commonly mutated oncogene TP53 (Benjamin  
123 & Ananthaswamy, 2006), which in normal function has an important role in tumor  
124 suppression. At low concentrations, reactive oxygen species (ROS) play vital roles  
125 during mutagenic activity in response to pathogen attack. Higher concentrations of ROS  
126 produced by UVR give rise to oxidative stress where ROS attack DNA bases and the  
127 deoxyribosyl backbone of DNA (see MacDavid & Aebisher, 2014). Production of  
128 antioxidant enzymes neutralizes ROS, and ROS modulation is controlled by the aryl  
129 hydrocarbon receptor (AhR) that plays a key role in mammalian evolution.

130  
131 The consequences of ionizing radiation associated with GCRs and solar particle events  
132 (solar wind) for human health have received attention in recent years in an effort to  
133 evaluate the health effects of future space travel outside Earth's protective  
134 magnetosphere (e.g., Delp et al., 2016). Earth's atmosphere is opaque to all but the  
135 highest energy GCRs, and, together with the geomagnetic field, serves to shield Earth's  
136 surface from GCRs. The intensity of UVR arriving at Earth's surface decreases with  
137 increasing latitude, and is attenuated by stratospheric ozone (O<sub>3</sub>) that acts as a sink for  
138 UVR. The geomagnetic field plays an important role in preserving the atmosphere,  
139 including stratospheric ozone that would otherwise be stripped away by solar wind and  
140 GCRs (Wei et al., 2014). UVR triggers dissociation of oxygen molecules (O<sub>2</sub>) into oxygen  
141 radicals that combine to form stratospheric ozone that in turn absorbs UVR as it splits  
142 into oxygen atoms. Certain ozone-depleting agents (such as nitrogen oxides) are

143 produced naturally by energetic particle precipitation (EPP) from solar wind,  
144 particularly during solar proton events (SPE), and therefore times of low geomagnetic  
145 field strength lead to higher ozone depletion (Randall et al., 2005, 2007). Atmospheric  
146 modeling implies substantial increases in hydrogen and nitrogen oxide concentrations  
147 due to enhanced ionization by GCRs during the Laschamp excursion, with significant  
148 decrease in stratospheric ozone particularly at high latitudes (Suter et al., 2014).  
149 Modeling of ozone depletion during polarity reversals, based on a geomagnetic field  
150 intensity  $\sim 10\%$  of the present value, leads to enhanced UVR flux at the Earth's surface,  
151 particularly at higher latitudes, that is 3-5 times that resulting from the anthropogenic  
152 ozone hole (Winkler et al., 2008; Glassmeier & Vogt, 2010). Prior to anthropogenic  
153 emission of ozone-depleting chlorofluorocarbons (CFCs) and halons, EPP at times of low  
154 geomagnetic field strength played an important role in ozone depletion. A well-defined  
155 nitrate peak, together with a broader  $^{10}\text{Be}$  peak, are associated with low field strength at  
156 the time of the Laschamp magnetic excursion ( $\sim 41$  ka) in the EPICA-Dome C Antarctic  
157 ice core (Traversi et al., 2014), which indicates that geomagnetic shielding played a role  
158 in the production of both cosmogenic isotopes (such as  $^{10}\text{Be}$ ) and ozone-depleting  
159 nitrogen compounds. It is noteworthy that bacterial UVR proxies in sediments from  
160 Lake Reid (Antarctica) imply more than three times higher UVR flux during part of the  
161 last glacial than during the Holocene (Hodgson et al., 2005). UVR exposure affects the  
162 early stage of life in modern marine plankton, and plankton-benthos coupling in coastal  
163 waters (e.g., Hernandez Moresino et al. 2011). Furthermore, UVR plays a role in  
164 photosynthesis (e.g., Hollosy, 2002) and can cause changes in vegetation, and habitat  
165 modification.

166  
167 The aim of this paper is to review the record of geomagnetic field intensity over the last  
168  $\sim 300$  kyr, and compare this record with the fossil record of extinction in terrestrial  
169 mammals (including Neanderthals), and with nodes in hominin phylogeny determined  
170 from mitochondrial DNA (mtDNA) and Y chromosomes.

171

## 172 **2. The geomagnetic field**

173

174 Knowledge of the Holocene geomagnetic field (i.e. over the last  $\sim 12$  kyr) has been based  
175 on models built from archaeomagnetic, lava and sediment data (Korte et al., 2011;  
176 Pavón-Carrasco et al., 2014; Constable et al., 2016). Beyond the Holocene, geomagnetic  
177 field strength during the Quaternary, over the last  $\sim 2.5$  Myr, has been acquired  
178 primarily from relative paleointensity (RPI) data from marine sediments recovered by  
179 deep-sea drilling (e.g., Laj et al., 2000, 2004; Valet et al., 2005; Ziegler et al., 2011;  
180 Channell et al., 2009, 2018; Xuan et al., 2016). Absolute paleointensity data from lavas  
181 are hampered by unknown time gaps between lava flows and inadequate age control in  
182 young (Quaternary) lavas. RPI data from sediments have been acquired by normalizing  
183 the natural remanent magnetization (NRM) intensity by a laboratory-acquired  
184 magnetization designed to activate the same population of magnetic grains that carry  
185 the NRM, thereby compensating for variations in concentration of NRM-carrying grains  
186 throughout the sedimentary section. The laboratory-applied normalizers are typically  
187 anhysteretic remanent magnetization (ARM) and/or isothermal remanent  
188 magnetization (IRM). ARM is acquired in a decreasing alternating field (AF, with a peak  
189 AF of  $\sim 100$  mT) with a weak direct current (DC) bias field (typically  $50 \mu\text{T}$ ), and IRM is  
190 acquired in a strong DC field (up to  $\sim 1$  T). The appropriate normalizer is usually chosen  
191 such that its coercivity (response to AF demagnetization) closely matches that of the

192 NRM. A typical RPI proxy comprises the slope of NRM intensity versus ARM (and/or  
193 IRM) intensity determined over a particular peak demagnetization field range (such as  
194 20-80 mT). The definition of the slope, often determined at 1-cm spacing down-core,  
195 can be gauged by determining its linear correlation coefficient (e.g., Xuan & Channell,  
196 2009).

197  
198 There is substantial agreement among late Quaternary stacks of global sedimentary RPI  
199 data (Valet et al., 2005; Channell et al., 2009; Ziegler et al., 2011; Xuan et al., 2016) that  
200 are independently supported by  $^{10}\text{Be}/^9\text{Be}$  data from marine sediments (e.g., Simon et al.,  
201 2016) due to the role of the geomagnetic field in modulating cosmogenic isotope ( $^{10}\text{Be}$ )  
202 production. Agreement is, however, poor among stacks of sedimentary RPI data and  
203  $^{10}\text{Be}/^9\text{Be}$  data in the 0-40 ka interval (see Fig. 1 of Channell et al., 2018) that is  
204 attributed to subtle and often unrecognized drilling disturbance in poorly consolidated  
205 uppermost sediments recovered from the ocean floor. A recently published RPI stack  
206 for 10-40 ka based on high-deposition-rate sediments from the Iberian Margin and  
207 elsewhere (Fig. 1) is supported by paleointensity estimates from revised calculation of  
208  $^{10}\text{Be}$  flux in Greenland ice cores (Channell et al., 2018) using the GICC05 Greenland ice-  
209 core age model (Svensson et al., 2008). Models and stacks covering the same 0-40 ka  
210 interval that use RPI data from lower sedimentation rate sequences (e.g., Panovska et  
211 al., 2018) cannot resolve the detail that is revealed by higher sedimentation rate  
212 sequences and by  $^{10}\text{Be}$  flux in ice cores.

213  
214 When adequately recorded, “magnetic excursions” are manifested in both lavas and  
215 sediments as antipodal magnetization directions defining short-lived polarity reversal,  
216 and they occupy minima in RPI records (Laj & Channell, 2015). At least five magnetic  
217 excursions have been documented in lavas and sediments of the last 300 kyr (Fig. 1)  
218 with ~13 in the Brunhes Chron (last 775 kyr). Directional magnetic excursions have  
219 been named after the location where they were initially recorded such as Mono Lake  
220 (34 ka), Laschamp (41 ka), Blake (94 ka and 120 ka), Iceland Basin (191 ka) and Pringle  
221 Falls (211 and/or 238 ka). Although the first record of a magnetic excursion (the  
222 Laschamp excursion) was published over 50 years ago (Bonhommet & Babkine, 1967),  
223 magnetic excursions remain controversial mainly because the aberrant magnetization  
224 directions that define them have sub-millennial to millennial duration and are,  
225 therefore, fortuitously recorded in geological archives. Although the Laschamp  
226 excursion has now been recorded in scores of globally distributed sedimentary  
227 sequences (e.g., Laj et al., 2000, 2006; Mazaud et al., 2002; Lund et al., 2005; Evans et al.,  
228 2007; Collins et al., 2012; Channell et al., 2017), the Laschamp records represent a tiny  
229 proportion of the total number of paleomagnetically-studied sedimentary sequences  
230 that cover the Laschamp interval. The sub-millennial duration of the excursion means  
231 that recordings of the excursion are generally restricted to sequences with mean  
232 sedimentation rates >10-15 cm/kyr that are unusual in deep-sea sediments (see  
233 Roberts & Winkelhofer, 2004). Lava records of the Laschamp are restricted to the  
234 Massif Central in France (e.g., Laj et al., 2014), the Auckland volcanic field (Cassata et al.,  
235 2008) and New Zealand’s Mt. Ruapehu (Ingham et al., 2017), that just happened to have  
236 erupted with high frequency during this brief period.

237  
238 The age of the Laschamp excursion, and the low geomagnetic field strength associated  
239 with it (Fig. 1), is based on correlation of deep-sea cores that record the Laschamp  
240 excursion to ice-core chronologies (e.g., Laj et al., 2000; Nowaczyk et al., 2012; Channell

241 et al., 2017), by  $^{40}\text{Ar}/^{39}\text{Ar}$  age determinations in lavas from the Massif Central (e.g., Laj  
 242 et al., 2014), by U/Th ages in a speleothem from Missouri (Lascu et al., 2016), and by the  
 243 age of the related cosmogenic-isotope flux peak in ice cores (Yiou et al., 1997;  
 244 Baumgartner et al., 1998; Wagner et al., 2000; Svensson et al., 2006, 2008; Traversi et  
 245 al., 2016). Current estimates of the duration and mid-point age of the Laschamp  
 246 directional excursion are <1 kyr and 41 ka, respectively. A less abrupt geomagnetic field  
 247 intensity decrease (over several kyr) defines a paleointensity minimum that brackets  
 248 the directional excursion.

249  
 250 Directional records of the Iceland Basin excursion at ~190 ka have been recovered from  
 251 numerous deep-sea sediment cores from the Northern Hemisphere (see review in  
 252 Channell, 2014), in cores from Lake Baikal (Oda et al., 2002), and two sites from the  
 253 South Atlantic ocean (Stoner et al., 2003; Channell et al., 2017). The age of the excursion  
 254 has been established by correlation to marine oxygen isotope records, and to ice core  
 255 records. There are no records of the Iceland Basin excursion from well-dated lava  
 256 sequences.

257  
 258 Apart from the Laschamp and Iceland Basin excursions, magnetic excursions during the  
 259 last 300 kyr are controversial because of the paucity of records and/or poor age control  
 260 in available records. On the other hand, RPI records and hence recordings of RPI  
 261 minima associated with the excursions are numerous, presumably because RPI minima  
 262 associated with excursions are manifest over longer timescales than the associated  
 263 directional excursions.

264  
 265 The Mono Lake excursion at ~34 ka has been recorded in both deep-sea sediments and  
 266 lava flows (Channell, 2006; Cassata et al., 2008; Kissel et al., 2011, Laj et al. 2014;  
 267 Negrini et al., 2014) as well as at Mono Lake in California (e.g., Benson et al., 2003),  
 268 although the age at the Mono Lake type-location is controversial (e.g., Kent et al., 2002;  
 269 Cassata et al., 2010). The age of a RPI minimum associated with the Mono Lake  
 270 excursion can be estimated from a cosmogenic-isotope flux maximum in Greenland ice  
 271 cores (Wagner et al., 2000; Muscheler et al., 2005).

272  
 273 Although the older of the two Blake excursions (94 and 120 ka) was first recorded  
 274 almost 60 years ago (Smith & Foster, 1969), there have only been a handful of  
 275 observations of either Blake excursion since (see Laj & Channell, 2015). The older Blake  
 276 excursion has, however, been observed in diverse media: marine sediments (e.g., Tric et  
 277 al., 1991), Chinese loess (e.g., Zhu et al., 1994) and a speleothem from Spain (Osete et al.,  
 278 2012). The younger “Blake” excursion, also known as the Skalamaelifell excursion, has  
 279 been observed in Icelandic lavas (Jicha et al., 2011). The age of the Pringle Falls  
 280 excursion (at 211 ka and/or 238 ka) is not yet settled due to a lack of records in well-  
 281 dated sequences (Herrero-Bervera et al., 1994; Singer et al., 2008; Singer, 2014; Laj &  
 282 Channell, 2015).

283  
 284 An apparent RPI minimum at ~13 ka appears in some individual RPI records, as a notch  
 285 in RPI stacks (Fig. 1), and in virtual axial dipole moment (VADM) proxies from  $^{10}\text{Be}$  flux  
 286 in Greenland ice cores (Channell et al., 2018). In one core from the Iberian Margin  
 287 (MD01-2444), a pronounced RPI minimum at ~13 ka is associated with a directional  
 288 magnetic excursion (Channell et al., 2013). The apparent RPI minimum at ~13 ka in  
 289 core MD01-2444 coincides with the old-end of the 0-14 ka Holocene paleointensity

290 model of Pavón-Carrasco et al. (2014). The RPI stack of Channell et al. (2018) contains  
 291 a prominent field intensity peak (VADM  $\sim 13 \times 10^{22}$  Am<sup>2</sup>) around 18-17 ka followed by a  
 292 field collapse between 16 and 14 ka culminating in a minimum VADM value ( $\sim 7 \times 10^{22}$   
 293 Am<sup>2</sup>) that has been smoothed by stacking (Fig. 1). The VADM decrease at 17-13 ka  
 294 corresponds to a rate of decrease (12.1 nT/yr) that is comparable to the rate (12.7  
 295 nT/yr) observed in the <sup>10</sup>Be-derived field at the same time (Channell et al., 2018), and to  
 296 the rates observed during the Laschamp/Mono Lake excursions (Laj & Kissel, 2015).

297

298 To provide a reference template for geomagnetic field intensity over the last  $\sim 300$  kyr,  
 299 we combine the RPI stack for 14-45 ka (Channell et al., 2018) with the Holocene model  
 300 for 0-14 ka (Pavón-Carrasco et al., 2014) and with the PISO paleointensity stack  
 301 (Channell et al., 2009) beyond 45 ka. This paleointensity template was calibrated to  
 302 VADM by aligning it with Holocene archaeomagnetic models (Korte et al., 2011; Pavón-  
 303 Carrasco et al., 2014; Constable et al., 2016) and assuming a value of  $\sim 1.5 \times 10^{22}$  Am<sup>2</sup> for  
 304 the VADM minimum at the Laschamp excursion (Laj et al., 2014). This VADM template  
 305 provides a proxy for geomagnetic intensity variations that is compared with events in  
 306 mammalian evolution, in order to test potential linkages between geomagnetic field  
 307 intensity variations and late Quaternary mammalian evolution.

308

### 309 **3. Late Quaternary Extinctions (LQE)**

310

311 Causes of extinction of mammalian megafauna (adult weight  $>45$  kg) during the Late  
 312 Quaternary, the so-called Late Quaternary Extinction (LQE), have been the subject of  
 313 prolonged debate (e.g., Martin, 1967; Koch & Barnosky 2006; Stuart 2015). Prior to  $\sim 13$   
 314 ka, the mammal assemblage of the Americas included large-bodied animals such as  
 315 mammoth, horses, camels, saber-tooth cats, and the short-faced bear. Extinction was  
 316 total for mammals larger than 1000 kg,  $>50\%$  for size classes between 32 and 1000 kg,  
 317 and  $\sim 20\%$  for those between 10 and 32 kg (Koch & Barnosky, 2006). Within a short  
 318 time window,  $>150$  species were lost in the Americas, including all mammals over  $\sim 600$   
 319 kg. An analogous catastrophic size-controlled LQE affected 14 of 16 Australian  
 320 mammalian genera; however, extinction ages are at least  $\sim 30$  kyr older than in North  
 321 America. Although fewer species were affected by the LQE in Africa and Eurasia, a  
 322 similar size-biased extinction has been observed with end-Pleistocene ( $\sim 13$  ka) ages  
 323 being prominent. Current explanations for the LQE in North America and Australia  
 324 involve a combination of two hypotheses: climate change, and “overkill” by human  
 325 hunting, modulated by the knock-on effect of herbivore extinction on the environment  
 326 and on the survivability of other groups (e.g., Owen-Smith, 1987). Although “overkill”  
 327 was originally used to explain North American extinctions (Martin, 1967), a forerunner  
 328 of the hypothesis was popular in 19<sup>th</sup> century Europe, before it was eventually  
 329 abandoned as archaeological evidence for human migration showed little evidence for  
 330 the impact of human hunting on the LQE. Grayson & Metzler (2003) argued that island  
 331 settings (e.g., New Zealand or the West Indies), where human hunting and habitat  
 332 degradation can be unequivocally associated with extinction, should not be the model  
 333 for continental extinctions. Extraterrestrial impact as a contributing cause for the LQE  
 334 in North America, and for the onset of the Younger Dryas cold period (Firestone et al.,  
 335 2007), have not been supported by subsequent analyses (Pinter et al., 2011; Holliday et  
 336 al., 2014).

337

338 In North America, the brief ( $\sim 200$  yr) duration of Clovis-tool finds at  $\sim 13$  ka (e.g.,

339 Waters & Stafford, 2007) is usually associated with rapid dispersal of modern humans  
340 across North America, and is closely contemporaneous with the LQE peak. On the other  
341 hand, the lack of mammalian kill sites in the Clovis record argues against a direct  
342 linkage between Clovis technology and “overkill”. Humans were present in North  
343 America at least several kyr prior to the Clovis horizon (Gilbert et al., 2008; Waters et  
344 al., 2011), and perhaps prior to ~24 ka (Bourgeon et al., 2017).

345

346 According to Faith & Surovell (2009), the LQE in North America was abrupt and  
347 requires a mechanism capable of wiping out ~35 genera across the continent in a  
348 “geological instant” in the 13.8-11.4 ka interval (Fig. 2), with the spread in last  
349 appearances being largely explained by an incomplete fossil record and the resulting  
350 Signor-Lipps effect (Signor & Lipps, 1982). Abrupt versus staggered megafaunal  
351 extinction at the LQE is central to determination of cause. In their “continental  
352 simulation”, Faith & Surovell (2009) determined the empirical probability (3.4 %) of  
353 observing a terminal Pleistocene (10-12 ka) age from 1955 stratigraphic occurrences  
354 (from 31 genera) of which 66 taxa (from 16 genera) yield terminal Pleistocene ages,  
355 assuming that all occurrences are equally likely to receive a terminal Pleistocene age.  
356 The simulation randomly assigned pre- or post- 12 ka ages to all 1955 observations  
357 based on this probability (3.4 %). The total number of genera that received a terminal  
358 Pleistocene age was tallied for each of 10,000 simulations to determine the probability  
359 of observing 16 or fewer terminal Pleistocene genera. The authors concluded that the  
360 observed pattern is consistent with synchronous (i.e., 10-12 ka) extinction for all 31  
361 genera.

362

363 Bradshaw et al. (2012) proposed a Gaussian-resampled, inverse-weighted McNerny  
364 (GRIWM) approach, which weights observations inversely according to their temporal  
365 distance from the last observation of a confirmed species occurrence, and samples  
366 radiometric ages from the underlying probability distribution. In Figure 2, we show  
367 GRIWM estimates of continent-wide European extinctions from the fossil record aided  
368 by DNA analyses (Cooper et al., 2015), excluding regional disappearances. An extinction  
369 age estimate in North America for *Arctodus simus* (the short-faced bear) at 10.8 ka  
370 (Schubert, 2010) and the onset of population decline of *Bison priscus* (the steppe bison)  
371 at ~37 ka (Shapiro et al., 2004) are included in Figure 2. Note that the horse and woolly  
372 mammoth (*Mammuthus primigenius*) persisted in interior Alaska until ~10.5 ka (Haile  
373 et al., 2009), and the woolly mammoth survived on St. Paul Island (Alaska) until ~5.6 ka  
374 (Graham et al., 2016). Zazula et al. (2014) pointed out that the American mastodon  
375 (*Mammuth americanum*) occupied eastern Beringia (Alaska/Yukon) during the last  
376 interglacial before its range contracted southward at the onset of glacial conditions at  
377 ~75 ka. The range of the species appears to have expanded northward again as  
378 interglacial conditions returned at the end of the Pleistocene, before extinction of the  
379 species at ~11.5 ka (10,000 <sup>14</sup>C years BP). Zazula et al. (2014) posed the question: why  
380 was this species stopped in its tracks when favorable conditions beckoned in Beringia?

381

382 An important proxy for herbivore population is the abundance in sedimentary  
383 sequences of coprophilous (dung) fungal spores, such as *Sporormiella*. The proxy was  
384 first proposed over 30 years ago (Davis, 1987), requires careful interpretation and  
385 laboratory techniques (e.g., van Asperen et al., 2016), but provides a measure of  
386 herbivore population independent of the bone-fossil record. Lake sediments in New  
387 York and Indiana imply a decline in *Sporormiella* spores beginning at 14.8 ka (Fig. 2)



388 that falls below the 2% threshold by 13.7 ka (Gill et al., 2009). This result has been  
389 closely replicated at Silver Lake (Ohio) where *Sporormiella* decline was dated at 13.9 ka  
390 (Gill et al., 2012). Importantly, the *Sporormiella* decline at these sites predates Younger  
391 Dryas cooling, and concurrent changes in the pollen record, and immediately predates a  
392 marked charcoal deposition increase, implying that herbivore decline and resulting  
393 landscape changes provide an explanation for subsequent (natural) landscape burning.  
394 The onset of the demise of North American herbivores at ~14.5 ka (Gill et al., 2009) lies  
395 within the Bølling-Allerød warm period with the Younger Dryas cold period beginning  
396 ~2 kyr later (e.g., Deplazes et al., 2013).

397  
398 The South American LQE was even more profound than that in North America (Koch &  
399 Barnosky, 2006; Barnosky & Lindsey, 2010), with the loss of 50 megafaunal genera  
400 (~83%). Robust dates are scarce for the South American LQE, although it appears that  
401 many taxa were lost near the Pleistocene-Holocene boundary (Barnosky & Lindsey,  
402 2010). *Sporormiella* decline in lake sediments from SE Brazil imply herbivore  
403 population collapse at ~12 ka (Raczka et al., 2018).

404 In Northern Eurasia, 9 genera (35%) were lost during the LQE. Available age data are  
405 consistent with a two-phase extinction in the 45-35 ka and 15-10 ka intervals (Koch &  
406 Barnosky, 2006). Up to 50% of worldwide megafaunal extinctions at 15-10 ka  
407 apparently occurred in Northern Eurasia (Cooper et al., 2015), but the extinction  
408 pattern is more complex than in North America with megafaunal range contractions  
409 culminating in extinction for some species but not others (Stuart, 2015). In Figure 2, we  
410 plot continent-wide Eurasian megafaunal extinction events from Cooper et al. (2015).  
411 Several well-studied species disappeared continent-wide at ~26-31 ka (Fig. 2) hence  
412 their last appearances significantly postdate the Laschamp excursion (at ~41 ka). Post-  
413 Laschamp extinction for *Crocota crocuta* (spotted hyaena) and *Crocota spelaea* (cave  
414 hyaena) at ~26 ka were, however, preceded by severe range contraction from Asia into  
415 Europe (Stuart & Lister, 2014). Fossils of *Ursus spelaeus* (cave bear) also indicate E to W  
416 range contraction before extinction at ~26 ka with abrupt population decline, based on  
417 DNA analyses, after 50 ka (Stiller et al., 2010; Stiller et al., 2014; Baca et al., 2016).

418 In Africa, at least 24 species and ~10 genera of mammals became extinct in the 13-6 ka  
419 interval, representing 25% of Pleistocene African megafauna (Faith, 2014). Species-  
420 level extinction was, again, most intense for larger-bodied megafauna (Koch &  
421 Barnosky, 2006). The African LQE was considered to have been less severe than  
422 elsewhere, accounting for the relatively rich diversity of modern African megafauna. On  
423 the other hand, the number of extinct African species dated to the last 100 kyr exceeds  
424 the number in Europe and matches the number in Australia, and is only surpassed by  
425 the LQE in the Americas (Faith, 2014). In east Africa alone, the number of securely  
426 dated latest Pleistocene mammal extinctions has risen from two to seven in the last  
427 decade, with most being of grazers associated with open habitats (Faith, 2014).

428 Estimation of extinction ages for Australian megafauna (and for some Eurasian genera)  
429 is complicated by the majority of last appearances being at or beyond the practical  
430 range of radiocarbon dating (i.e., >40 ka). The LQE in Australia was apparently  
431 catastrophic for large mammalian megafauna, with the complete loss of all animals  
432 heavier than ~100 kg. Fourteen of sixteen genera of Pleistocene mammalian megafauna  
433 disappeared, together with all megafaunal reptiles (6 genera), in the vicinity of, or prior

434 to, ~40 ka (Fig. 2). Ten Australian genera disappeared in the 44-35 ka interval based on  
 435 a variety of frequentist statistical methods (including GRIWM) to determine extinction  
 436 ages for 16 megafaunal genera (Fig. 2; Saltré et al., 2016). The mass extinction of  
 437 megafauna at this time, including the largest-known (~3000 kg) marsupial  
 438 (*Diprotodon*), has been linked with climate variability and aridity (e.g., Wroe et al.,  
 439 2013) although this linkage has been disputed (e.g., Saltré et al. 2016), often in favor of  
 440 human predation or “overkill” (e.g., Brook & Johnson, 2006; Miller et al., 2016; Johnson  
 441 et al., 2016; van der Kaars et al., 2017). It is noteworthy that the extinction age for the  
 442 ~200-kg flightless bird *Genyornis newtoni* at ~35 ka (Fig. 2; Saltré et al., 2016) is  
 443 younger than the ~43 ka estimate given by Miller et al. (2016) based on dated eggshell  
 444 fragments. Even if final extinction was delayed until ~35 ka, the population of *Genyornis*  
 445 *newtoni* crashed close to the time of the Laschamp excursion (~41 ka) based on the egg-  
 446 shell data (Miller et al., 2016), although egg-shells attributed to *Genyornis newtoni* may  
 447 be from other species (Grellet-Tinner et al., 2016).

448  
 449 At Lynch’s Crater (NE Queensland), an abrupt decline in dung fungi including  
 450 *Sporormiella* (Figs. 2 and 3) implies abrupt demise of large Australian herbivores at 40-  
 451 44 ka (Johnson et al., 2015). An abrupt increase in charcoal lags *Sporormiella* decline by  
 452 ~100 years, and evidence for grasses and sclerophyll vegetation lags *Sporormiella*  
 453 decline by ~300-400 years (Rule et al., 2012; Johnson et al., 2015). The charcoal-rich  
 454 levels can be explained by natural lightning-induced biomass burning as a result of fuel  
 455 build-up triggered by herbivore extinction (Rule et al., 2012; Johnson et al., 2016). Off  
 456 the southern coast of Western Australia, marine core MD03-2614G records a sharp  
 457 decline in *Sporormiella* in the 45-43 ka interval, relative to values recorded back to 140  
 458 ka (Figs. 2 and 3; van der Kaars et al., 2017).

459  
 460 A role for humans in the extinction of large animals in Australia remains popular (e.g.,  
 461 Brook & Johnson, 2006; Turney et al., 2008; Miller et al., 2016; Johnson et al., 2016; van  
 462 der Kaars et al., 2017), although the arrival of humans in Australia (Sahul) may have  
 463 predated the LQE at ~40 ka by ~25 kyr (Clarkson et al., 2017) although the arrival date  
 464 is not unequivocal (O’Connell et al., 2018). There is no evidence for a spike in the human  
 465 population in Australia at the time of the most prominent extinction event at ~40 ka,  
 466 when the entire Australian human population may not have exceeded a few tens of  
 467 thousands (Williams, 2013). Tasmania had a land bridge to the Australian mainland  
 468 during the last glacial, becoming an island in the early Holocene. The extinction of  
 469 megafauna in Tasmania at ~40 ka does not correspond to climate or environmental  
 470 change, and has been associated with the late arrival of humans in the region (Turney et  
 471 al., 2008). More recent results place the Tasmanian extinction of *Protemnodon anak* and  
 472 other megafauna at ~41 ka, predating human arrival on the island at ~39 ka and hence  
 473 precluding human involvement in the extinctions (Cosgrove et al., 2010; Lima-Ribeiro &  
 474 Diniz-Filho, 2014). Extant smaller (more accessible) prey, particularly the common  
 475 wombat (*Vombatus ursinus*) and the red-necked wallaby (*Macropus rufogriseus*),  
 476 characterize the early archaeological kill-sites on the island (Cosgrove et al., 2010).

477  
 478 Apart from Australian extinctions concentrated close to the Laschamp excursion at ~40  
 479 ka, fossils from the King’s Creek Catchment (SE Queensland) indicate additional  
 480 concentrations of megafaunal last appearances at ~83 ka, ~107 ka and ~122 ka (Price  
 481 et al., 2011; Wroe et al., 2013). The older two dates (107 ka and 122 ka) correspond to  
 482 magnetic field intensity minima associated with the Blake excursions (Fig. 1b).

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#### 4. Neanderthal extinction

The extinction of Neanderthal represents one of the great puzzles of human evolution. Neanderthal and anatomically modern humans (AMHs) cohabited Western Europe for ~2-5 kyr, prior to ~39 ka, supporting the contention that competition may have contributed to the demise of Neanderthal (Higham et al., 2014). Brief cold and dry conditions in Europe associated with Heinrich Stadial (HS) 4 were proposed as an additional likely stressor on Neanderthal (Sepulchre et al., 2007). Analyses of Campanian Ignimbrite (CI) cryptotephra from archaeological sites in Greece and elsewhere in Eastern Europe and Libya indicated that the CI eruption occurred early in a dry period associated with HS4, postdated the end of the Middle Paleolithic and the Mousterian tool industry, and hence postdated the demise of Neanderthal (Lowe et al., 2012). In Black Sea sediment cores, CI tephra overlie, and therefore postdate, the Laschamp excursion (Nowaczyk et al, 2012).

The extinction of Neanderthal and the demise of the Mousterian tool industry (Fig. 3) can be placed at 41,030–39,260 calibrated years before present (41-39 ka) with 95.4% probability (Higham et al., 2014). Cooper et al. (2015) estimated the extinction of Neanderthal at 41,227 calibrated years before present (BP) with a standard deviation of 219 years, and 39,528-41,013 calibrated years BP using the GRIWM method (Fig. 2). Mousterian ages outside this range have been recorded at several locations in southern Iberia including Gorham’s Cave in Gibraltar (Finlayson et al., 2006; Tzedakis et al., 2007), but these ages should now be disregarded according to Higham et al. (2014). Recent findings cast doubt on the existence of Neanderthal after ~39 ka, and lead to a closer correspondence of the demise of Neanderthals with the Laschamp magnetic excursion and the associated brief interval of very low geomagnetic field intensity centered at ~41 ka (Laj et al., 2014). It is important to note that the IntCal13 radiocarbon calibration (Reimer et al., 2013) may be offset to older ages by ~1 kyr in the vicinity of the Laschamp excursion, relative to ice-core chronologies (Muscheler et al., 2014).

Valet & Valladas (2010) proposed that low magnetic field strength in the Laschamp/Mono Lake excursion interval (40-33 ka) was an important factor in Neanderthal demise. Why anatomically modern humans (AMHs) were not similarly affected has remained an open question considering that the two populations shared habitats for 2600-5400 years (Higham et al. 2014) or >5 kyr (Lowe et al., 2012). There is no evidence for differences in skin pigmentation between European AMHs and Neanderthals, and at least a fraction of Neanderthals apparently had the same pale skin and/or red hair observed in some modern humans (Lalueza-Fox et al., 2007). Natural skin pigmentation in humans mitigates the harmful effects of UVR but its advantage is offset by the importance of sunlight for vitamin D3 synthesis. The skin protection factor (SPF) of “red ochre” (hematitic iron oxides) is traditionally utilized by some African tribes (Rifkin et al., 2015) and has been used since at least the last interglacial (~120 ka) based on ochre coatings on strung beads and residues on storage shells from Africa (Hodgskiss & Wadley, 2017), SE Spain (Hoffmann et al., 2018) and Levantine Mousterian sites (Bar-Yosef Mayer et al., 2009). The mystery of AMH survival at the time of Neanderthal demise may have been resolved by the discovery of differences in amino acid substitution in an intracellular chemosensor (the aryl hydrocarbon receptor,

532 AhR) for AMHs and for Neanderthals and other primates (Hubbard et al., 2016).

533

## 534 **5. The role of the aryl hydrocarbon receptor (AhR)**

535

536 Defense mechanisms against UVR include the production of quenching agents and anti-  
537 oxidant enzymes that neutralize reactive oxygen species (ROS) produced by UVR. The  
538 ROS modulation is controlled by the AhR intracellular chemosensor that plays a key role  
539 in the evolutionary response to UVR. Experimental results indicate an adaptive  
540 response of mitochondria to varying ROS levels under a phenomenon called  
541 mitohormesis (Becker et al., 2016).

542

543 The primary role of AhR is to regulate the transcription of genes mediating responses to  
544 the biochemical and toxic effects of dioxins, polyaromatic hydrocarbons, and related  
545 compounds (Abel & Haarmann-Stemann, 2010). AhR is expressed in all skin cells and  
546 can be generated by UVR through an endogenous ligand formed in situ from an amino  
547 acid called tryptophan (Esser et al. 2009). UVR and the more harmful UVB (wavelength  
548 290-320 nm) is absorbed by free tryptophan in the cytosol of epidermal cells, and AhR  
549 plays a key role in translocating UVR stress response to the nucleus (Wei et al., 1999;  
550 Fritsche et al., 2007; Tigges et al., 2014). Exposure to UVR, particularly UVB, generates  
551 highly mutagenic DNA photoproducts. The process initiates apoptosis and involves  
552 damage to nuclear DNA accompanied by mitochondrial dysfunction (Frauenstein et al.,  
553 2013). There is a general consensus that the AhR of modern humans is implicated in  
554 DNA repair (Schreck et al., 2009; Dittmann et al., 2016), tumor suppression (Fan et al.,  
555 2010; Yu et al., 2017), epidermal barrier function (Noakes 2015), skin tanning response,  
556 and melanocyte homeostasis (Luecke et al., 2010; Jux et al., 2011). Phylogenetic analysis  
557 suggested that the ability of vertebrate AhR to sense xenobiotics was acquired at a late  
558 stage of evolution, implying that the driving force for evolutionary conservation of AhR  
559 lies not only in its role in xenobiotic metabolism but also in normal cell development  
560 (Hao & Whitelaw, 2013; Hahn et al., 2017).

561

562 Hubbard et al. (2016) showed that the AhR variant in modern humans contains Val381  
563 residue in the ligand-binding domain, while the AhR of Neanderthals, and a Denisovan  
564 individual, as well as non-human primates and other vertebrates (rodents) encode the  
565 ancestral Ala381 variant. The Val381 variant is fixed in the genome of all modern  
566 humans as well as in the genome of the oldest (45 ka) AMH individual sequenced to  
567 date (Fu et al., 2014). Hubbard et al. (2016) suggested that the unique modification of  
568 AhR in AMHs led to significant competitive advantage over their Neanderthal neighbors,  
569 due to decreased sensitivity in AMHs to toxins associated with fire-smoke, the effects of  
570 which may have been exacerbated by troglodytic lifestyles.

571

572 Our focus here is on AhR involvement in the regulation of the skin responses to UVR,  
573 especially to harmful UVB radiation, and its modulation of the immune system (Rannug  
574 & Fritsche, 2006; Agostinis et al., 2007; Esser et al., 2013). UVB induces two signaling  
575 routes in mammalian cells: first, UVB is absorbed by nuclear DNA that results in  
576 generation of DNA photoproducts, and second, UVB activates cell-surface receptors  
577 (Merk et al., 2004). AhR plays an important role in skin integrity and immunity. AhR  
578 activation leads to transcriptional gene activation, and is involved in the cutaneous  
579 stress response to UVR (Agostinis et al., 2007; Dittmann et al., 2016; Schwarz, 2005;  
580 Navid et al., 2013) and alterations of gene expression (Dugo et al., 2012). Activation of

581 AhR by UVB leads to signaling both to the nucleus and to cell membranes (Fritsche et al.,  
 582 2007). The findings show that UVB irradiation affects cell surface receptors with  
 583 subsequent activation of mitogen-activated protein kinases that in turn affect DNA in  
 584 the nucleus.

585

## 586 **6. The early fossil record of *Homo sapiens***

587

588 From analyses of the fossil and stone-tool record over the last 250 kyr (Fig. 4), Lahr  
 589 (2016) proposed five transitions in the evolutionary history of *Homo sapiens*. (1) The  
 590 origins of the species at 240–200 ka; (2) the first major expansion at 130–100 ka; (3) a  
 591 period of dispersals at 70–50 ka; (4) a period of local/regional structuring of diversity  
 592 at 45–25 ka; and (5) an early Holocene phase of significant extinction of hunter-  
 593 gatherers and expansion of farmers (the Holocene Filter).

594

595 Until recently, Member 1 of the (Omo) Kibish Formation of Ethiopia yielded the earliest  
 596 known AMH cranial remains (Day et al., 1969). A volcanic tuff about 2 m below the level  
 597 of the fossil finds has a  $^{40}\text{Ar}/^{39}\text{Ar}$  weighted mean age of  $196 \pm 2$  ka (McDougall et al.,  
 598 2005). An age of 196 ka for Member 1, combined with sedimentological evidence for  
 599 rapid deposition, are consistent with Member 1 having been deposited synchronously  
 600 with Mediterranean sapropel S7 (McDougall et al., 2008, Brown et al. 2012). Jebel  
 601 Irhoud, Morocco, has been an important archaeological site since the 1960s when  
 602 human fossils were found alongside Mousterian stone tools and were once dated at ~40  
 603 ka (Ennouchi, 1962). Recent fossil discoveries at this location support the presence of  
 604 fossils with characteristics of AMHs (Hublin et al., 2017), and new luminescence age  
 605 dating (Richter et al., 2017) indicates that the fossils are considerably older, at ~300 ka,  
 606 than the Ethiopian AMH finds at Omo Kibish (Fig. 4). The finds in Morocco are now  
 607 among the earliest known hominin fossils with AMH characteristics, in common with  
 608 characteristics of the Florisbad fossil from South Africa discovered by T.F. Dreyer in  
 609 1932 and dated at ~260 ka (Grün et al., 1996). With such a sparse hominin fossil record,  
 610 it is not possible to predict when AMHs first appeared, although the period from 300 ka  
 611 to 200 ka in Africa appears to have been a critical time in development of AMH, and may  
 612 constitute the evolutionary cradle of our species.

613

## 614 **7. Time to most recent common ancestor (TMRCA) from mtDNA and Y- 615 chromosomes**

616

617 Thirty years ago, Cann et al. (1987) demonstrated that ethnically diverse surveys of  
 618 modern mitochondrial DNA (mtDNA) are a major source of human evolutionary history.  
 619 The advantages of mtDNA as an evolutionary tool include faster mutation rates in  
 620 mtDNA than in nuclear genes, and mtDNA is inherited maternally and does not  
 621 recombine. Cann et al. (1987) demonstrated the African origin of the human  
 622 mitochondrial gene pool, and estimated that mtDNA stems from an African  
 623 mitochondrial “Eve” who lived approximately 200,000 years ago. Initially, this  
 624 conclusion met with considerable resistance, however, increasing numbers of studies,  
 625 including work on Y-chromosomes, imply that the “Eve” hypothesis is substantially  
 626 correct. There is now broad consensus for the “out-of-Africa” hypothesis whereby  
 627 modern humans appeared at ~200 ka in Africa and spread throughout the continent  
 628 before dispersing across the globe, although the exact chronology and nature of the  
 629 population divergence remains unclear (Zhou & Teo, 2016). Dispersal has resulted in

630 the occupation of a wide variety of habitats with selection in response to specific  
631 ecological pressures. A complete understanding of adaptation depends on a description  
632 of the genetic mechanisms and selective history that underlies heritable traits (e.g.,  
633 Radwan & Babik, 2012), and the signatures of natural selection are a response to  
634 selective pressures that are often unknown. Estimating the age of selection signals may  
635 allow reconstruction of the history of environmental changes that shaped human  
636 phenotypes with specific ages associated with human dispersal out-of-Africa and the  
637 spread of agriculture (Nakagome et al., 2016).

638  
639 Both mtDNA and Y-chromosomes have been used to reconstruct human history. The  
640 former (mtDNA) is often the better option for analyzing ancient DNA because it is easier  
641 to obtain, is present in higher number in human cells, and does not undergo  
642 recombination. However, mtDNA reflects only the maternal history of a population and  
643 the history of a single individual may not accurately reflect the history of a population.  
644 For this reason, mtDNA studies should be complemented by data on the male-specific Y-  
645 chromosome (Pakendorf & Stoneking, 2005). Rates of human evolution have been  
646 estimated by applying a number of substitution models to mtDNA or Y-chromosome  
647 sequence data (Behar et al., 2008; Soares et al., 2013; Wang et al., 2014; Kivisild, 2015).  
648 In some examples, calibration has relied on the assumption that the genetic separation  
649 between humans and chimpanzees occurred at ~6 Ma, and that the evolutionary  
650 process has been clocklike since that time, while more complex models involve different  
651 substitution rates for coding and control regions of the genome (see Endicott et al.,  
652 2009). Molecular clocks can be tested by comparison with archaeological data  
653 pertaining to human migration.

654  
655 Phylogeny based on the first complete mtDNA sequence data available in  
656 year 2000 (from 52 individuals selected from around the world) yielded a  
657 time to the most recent common ancestor (TMRCA) of ~190 ka (see  
658 Oppenheimer, 2003, 2009; Soares et al. 2009), apparently consistent with the  
659 conclusions of Cann et al. (1987) and with dated fossil finds of early AMH in Ethiopia  
660 (McDougall et al., 2005). More recent studies of human evolution from mtDNA  
661 (Oppenheimer, 2012), and of paternal evolution from Y-chromosomes (Poznik et al.,  
662 2016), yield a broadly consistent picture of human evolution over the last ~200 kyr  
663 although the timing of branches in the evolutionary tree are poorly constrained and  
664 depend on estimated rates of mutation and population size (Fig. 4). As pointed out by  
665 Wang et al. (2014), Y-chromosomal substitution rates obtained using different  
666 calibration modes vary considerably, and produce disparate reconstructions of human  
667 history. An additional determinant of substitution rate is the efficacy of purifying  
668 selection, which in turn depends not only on the particular constraints of each  
669 chromosome, but also on the long-term effective population size for each chromosome  
670 type (Elhaik et al., 2014). Furthermore, mutation rates may have changed during  
671 hominoid evolution (Scally, 2016).

672  
673 The TMRCA is not a unique number, but rather a probability distribution based on two  
674 fundamental assumptions: the number of mutations and the mutation rate. Estimates of  
675 TMRCA depend strongly on the substitution rate and different results can be obtained  
676 by using different rates. For example, Mendez et al. (2013) estimated a very early date  
677 of 338 ka for the TMRCA of the Y-chromosome tree (L00) from a population of African-  
678 Americans. The authors explained this early age by either long-standing population

679 structure among modern human populations or archaic introgression from unknown  
680 species into the ancestors of modern humans in western Central Africa. However, other  
681 researchers (e.g., Wang et al., 2014) have pointed out that this ancient TMRCA can be  
682 partially attributed to the low substitution rate used by the authors. By using either a  
683 higher mutation rate or more extensive sequencing data, the estimate of TMRCA  
684 becomes much younger at about 208 ka (Elhaik et al. 2014) or 257 ka (D'Atanasio et al.  
685 2018).

686  
687 Poznik et al. (2013) reported the entire Y-chromosome and mitochondrial genome  
688 sequences using a within-human calibration point to estimate the substitution rate. The  
689 results indicate small TMRCA differences from Y-chromosomes (120–156 ka) and  
690 mtDNA (99–148 ka) that disagree with the conventional suggestion that the common  
691 ancestor of male lineages lived considerably more recently than that of female lineages.  
692 Analyzing the entire Y-chromosome dataset of the 1000 Genomes Project (2012, 2015),  
693 using the pedigree-based substitution rate, Wang et al. (2014) estimated the TMRCA at  
694 105 ka that is consistent with the estimate (105 ka) of Cruciani et al. (2011) and the  
695 estimate (101–115 ka) of Wei et al. (2013).

696  
697 Phylogeny based on mtDNA sequence data has provided an estimate of TMRCA (190 ka)  
698 with branching episodes at ~120 ka, ~70 ka, ~40 ka and ~15 ka which coincide with  
699 AMH dispersal patterns (Oppenheimer, 2012). Wei et al. (2013) used five substitution  
700 models to assess phylogenetic nodes from the TMRCA of 29 Y-chromosomes that  
701 yielded branching ages for one of the models (GENETREE-2) of  $112 \pm 12$  ka,  $68 \pm 7$  ka,  $49$   
702  $\pm 6$  ka and  $13 \pm 2$  ka. Applying the BEAST method to 68 worldwide Y-chromosomes,  
703 Scozzari et al. (2014) dated the first two splits in their tree at ~196 ka and ~167 ka,  
704 followed by TMRCA at ~110 ka, 85–77 ka, 51–33 ka and 8–22 ka (Fig. 4).

705  
706 It remains challenging to reconstruct population structure prior to ~60 ka using  
707 existing data, but the modern-human African mtDNA pool contains phylogenetic  
708 patterns that can be used to estimate the ages of several haplogroups. Analyzing L0  
709 HVS-I sequences, Rito et al. (2013) classified five branches (L0a, L0b, L0d, L0f, and L0k)  
710 of the mtDNA tree. The age estimates indicate that the mtDNA tree split to form L0 at  
711 ~180 ka, and later diversity follows a geographical distribution from southern Africa  
712 northward. Around ~128 ka, two distinct AMH groups co-existed in Africa with a first  
713 split around 119 ka (L0k) followed by major clades at 98.7 ka (L0f), 70.9 ka (L0b) and  
714 42.4 ka (L0a). It is noteworthy that the splitting of the widespread and common L0a, the  
715 diversification of L0k, and population increases of both haplogroups L0 and L3 can be  
716 dated to ~40 ka (Fig. 4).

## 717 718 **8. Discussion**

719 Past episodes of low geomagnetic dipole field strength affect UVR levels arriving at the  
720 Earth's surface because reduction in magnetospheric shielding results in lower  
721 stratospheric ozone levels, and hence lowered UVR shielding (Wei et al., 2014). Solar  
722 storms enhance nitrogen oxide production in the stratosphere that, in early 2004, led to  
723 ozone reduction of more than 60% at high northern latitudes (Randall et al., 2005).  
724 Similarly, atmospheric N<sub>2</sub>O concentrations are enhanced during times of low  
725 geomagnetic field strength when shielding from solar storms and GCRs is diminished.  
726 Earth's surface naturally emits N<sub>2</sub>O from the oceans and from soils, with emissions

727 having increased due to anthropogenic practices. N<sub>2</sub>O emissions are enhanced during  
728 interglacial and interstadial (warm) climate states, and increased by ~50% at the last  
729 glacial termination (Schilt et al., 2013, 2014). Enhanced atmospheric N<sub>2</sub>O  
730 concentrations during the Bølling-Allerød warm period, which coincided with an  
731 apparent geomagnetic field strength minimum at ~13 ka, would have elevated UVR  
732 reaching the Earth's surface at this time. It is noteworthy that the drastic magnetic field  
733 intensity decrease after the peak at 18 ka (Fig. 1) coincided with paleoclimatic changes  
734 especially in the southern hemisphere (Boex et al., 2013; Moreno et al., 2015; Martinez-  
735 Garcia et al., 2014). Evidence from Antarctic ice cores indicates sudden enhanced  
736 tropospheric UVR that has been related to a Mount Tahahe eruption dated to 17.7 ka  
737 (McConnell et al., 2017).

738

739 A selection process resulted Neanderthal disappearance at ~41 ka that apparently did  
740 not affect AMHs. Previous hypotheses for Neanderthal disappearance, and expansion of  
741 AMHs, include differences in subsistence strategies, language skills, and technical,  
742 economic and social systems, and the ability to adapt to changing environments (e.g.,  
743 Kochiyama et al., 2018). Response to environmental change such as UVR flux at Earth's  
744 surface would have involved the AhR, a chemosensor that regulates immunity and  
745 differs in AMHs versus Neanderthals and other primates (Hubbard et al., 2016). Ages  
746 for the end of the Mousterian tool industry and Neanderthal demise (Higham et al.,  
747 2014) are now tightly constrained to the Laschamp magnetic excursion (at 41 ka)  
748 implying a role for high UVR levels during the Laschamp field intensity minimum (Fig.  
749 3).

750

751 Prominent low geomagnetic field intensity episodes at 285 ka, 190 ka, 110-120 ka, 64  
752 ka, 41 ka and 13 ka (Fig. 1) appear to correspond to important times in evolution of  
753 hominins and other large mammals. At the LQE, megafauna in Australia, Europe and the  
754 Americas were thought to have become extinct over a protracted time in the Late  
755 Quaternary, however, improved age estimates and discovery of new fossil sites (Roberts  
756 et al., 2001; Koch & Barnosky, 2006; Faith & Surovell, 2009; Barnosky & Lindsey, 2010;  
757 Price et al., 2011; Wroe et al., 2013; Stuart, 2015; Faith, 2014; Cooper et al., 2015; Saltré  
758 et al., 2016) have led to extinction peaks becoming progressively constrained to ~13 ka  
759 in the Americas, to ~40 ka (with earlier episodes at 84, 107 and 122 ka) in Australia,  
760 and a complex combination of late Pleistocene ages (including ~13 ka and ~40 ka) in  
761 Eurasia and Africa.

762

763 One of the outstanding and intriguing aspects of the LQE is the strong correlation  
764 between extinction and body mass. As the vast majority of cell mutations are  
765 deleterious, large long-lived organisms are at an evolutionary disadvantage. In addition,  
766 small mammals often have opportunities to avoid UVR through burrowing. In modern  
767 mammal populations, there is, however, no apparent correlation between body mass  
768 and cancer occurrence, known as Peto's Paradox (Peto et al., 1975). The elephant  
769 genome includes 20 copies of an oncogene (TP53) that is a crucial tumor suppressor  
770 gene involved in apoptosis in response to DNA damage, whereas other mammals  
771 usually have small numbers of this gene (Abeggelen et al., 2015; Sulak et al., 2016). This  
772 discovery may explain why elephants are one of few large mammals to pass the LQE  
773 barrier. On the other hand, the remains of two extinct mammoth species included more  
774 than a dozen copies of TP53 in their genomes, and the American mastodon, that  
775 disappeared at ~ 50 ka, had 3-8 copies in its genome which implies an evolutionary



776 selective trend in the increased number of copies of TP53 in the genomes of these  
777 related megafauna (Sulak et al., 2016). TP53 is considered the guardian of the genome  
778 due to its role in mitigating DNA damage, and is itself a target of UV-induced mutations  
779 (Aylon & Oren, 2011; de Pedro et al. 2018).

780  
781 Huang et al. (2017) analyzed the body weight of species in two orders of large ungulate  
782 herbivores (Artiodactyla and Perissodactyla) from the Neogene fossil record in Europe  
783 and North America. They found a significant and progressive increase in body weight  
784 from early Miocene to late Pliocene for both orders in North America and for  
785 Artiodactyla in Europe. This work was followed by the analysis of Smith et al. (2018)  
786 who documented a global increase in body-size of megafauna during the Cenozoic, with  
787 abrupt downsizing at the LQE in late Pleistocene. We speculate that large mammals may  
788 have reached a natural body-size limit by late Pleistocene, due to increased likelihood of  
789 cell mutation at times of high UVR flux, as life spans and body weights increased during  
790 the Cenozoic. The envisaged role of UVB in reducing megafaunal populations at the LQE  
791 does not involve an instantaneous “blitzkrieg”, but rather an accumulation of UVB-  
792 triggered mutations over multiple (~30) generations or ~0.5-1 kyr, the approximate  
793 duration of RPI minima associated with magnetic excursions.

794  
795 There is no clear correlation of LQE events, either at ~13 ka (in North America and  
796 Eurasia), or ~40 ka (in Australia and Eurasia), with the first appearances of humans,  
797 which apparently preceded the LQE by at least 10 kyr in Australia, and at least several  
798 kyr in North America. The human population of Australia at ~40 ka was likely no more  
799 than a few tens of thousands, with no evidence of an increase in population at this time  
800 (Williams, 2013). According to Webb (2013), the “overkill” hypothesis “*is more*  
801 *sensational, granted, but the arguments are unrealistic and the evidence for it, at least in*  
802 *Australia, is non-existent*”. The case for “overkill” in North America at ~13 ka involves  
803 the close coincidence of the LQE with the brief (few century duration) Clovis-tool  
804 horizon. On the other hand, based on *Sporormiella* decline in lake sediments from the  
805 eastern USA, the demise at North American herbivores began at ~14.5 ka (Gill et al.,  
806 2009, 2012), prior to the first appearance of Clovis tools. Clovis tools have only rarely  
807 been found in association with megafaunal remains. Evidence for pre-Clovis human  
808 occupation in the Americas includes locations in southern Chile (Dillehay et al., 2015),  
809 and the Florida panhandle where human occupation at ~14.5 ka predated the  
810 *Sporormiella* decline by ~2 kyr (Halligan et al., 2016), consistent with *Sporormiella*  
811 disappearance at ~12.7 ka at another Florida panhandle location (Perotti, 2018).  
812 Paucity of evidence for human occupation prior to ~13 ka can be attributed to  
813 habitation being concentrated in coastal regions that were largely submerged during  
814 the subsequent (last) deglaciation, when sea-level rise necessitated migration of  
815 humans and other terrestrial mammals into the continental interior. We speculate that  
816 large mammals, particularly those that did not burrow, were particularly susceptible to  
817 DNA damage associated with low geomagnetic field strength and the resulting increase  
818 in UVB reaching the Earth’s surface. The coincidence of low magnetic field strength with  
819 LQE events, particularly at ~40 ka and ~13 ka, but also at 107 ka and 122 ka in  
820 Australia, implies that UVB flux was a contributing cause of the LQE in North America,  
821 Europe and Australia.

822  
823 The ~190 ka paleointensity minimum at the Iceland Basin excursion corresponds to  
824 TMRCA determined from the mtDNA and Y-chromosomes of modern humans (Gonder

825 et al., 2007; Soares et al. 2009; Oppenheimer, 2009, 2012; Wei et al., 2013; Poznik et al.,  
 826 2016). Other paleointensity minima at 110-120 ka, 64 ka, 41 ka and 13 ka (Fig. 1)  
 827 correspond to branches in phylogeny estimated from mtDNA and Y-chromosome  
 828 analyses (Fig. 4). For example, Wei et al. (2013) used five models to assess phylogenetic  
 829 nodes from the TMRCA of 29 Y-chromosomes that yielded branching ages for one model  
 830 (GENETREE-2) of  $112 \pm 12$  ka,  $68 \pm 7$  ka,  $49 \pm 6$  ka and  $13 \pm 2$  ka, corresponding closely to  
 831 minima in relative paleointensity records (Figs. 1 and 4). Although the transition from  
 832 Middle Paleolithic to Upper Paleolithic at  $\sim 40$  ka did not correspond to significant  
 833 changes in AMH anatomy in the fossil record, it is an important time for human  
 834 population structure (Fig. 4; Lahr, 2016) with mtDNA implying rapid population growth  
 835 at a time of major advance in tool design and the first appearance of artwork  
 836 (Stoneking, 1994).

837  
 838 Phylogeography, the study of human dispersal, demonstrates that when migration  
 839 occurs from one region to another, new mutations unique to that region accumulate  
 840 (Soares et al. 2016). Local adaptation to different habitats, including changes in exposure  
 841 to mutagenic solar radiation, partially controlled by the magnetic field, are potential  
 842 sources of phenotypical divergence (Jablonski & Chaplin, 2000, 2010). During  
 843 migrations, humans have adapted to differences in climate, altitude, and resource  
 844 availability. Migration to new environments alters selection pressures on the human  
 845 genome, and genetic studies have identified certain loci that were likely targets of this  
 846 selection. For example, highly pigmented skin protects against skin cancer but reduces  
 847 synthesis of vitamin D3, so differences in the amount of UVR place differing selection  
 848 pressures on pigmentation genes (Creanza & Feldman, 2016).

## 849 **9. Conclusions and Outlook**

851  
 852 Although coincidence does not prove causality, the timing of geomagnetic field strength  
 853 minima (hence enhanced UVR flux at Earth's surface) appears to correspond to events  
 854 in mammalian evolution. Improvements in knowledge of past geomagnetic field  
 855 strength, new mammalian fossil finds, advances in radiocarbon dating and DNA  
 856 analyses of fossils, use of dung-fungal proxies for herbivore populations, and advances  
 857 in the use of mtDNA and Y-chromosomes to map human phylogeny, have all contributed  
 858 to this possible linkage. Minima in geomagnetic dipole field strength at  $\sim 13$  ka and  $\sim 41$   
 859 ka (Laschamp, Fig. 1) led to stratospheric ozone depletion and UVB levels at the Earth's  
 860 surface that may have contributed to extinction of large mammals at these times,  
 861 although extinction dates are associated with large errors due to the inadequacy of the  
 862 fossil record (Fig. 2). The Neanderthals were apparently victims of the  $\sim 41$  ka  
 863 (Laschamp) magnetic field minimum, and differences in the AhR of modern humans and  
 864 Neanderthals may explain why we passed the Laschamp evolutionary barrier. Episodes  
 865 of low field strength further back in time notably at  $\sim 64$  ka,  $\sim 110$ -120 ka (Blake, Fig. 1),  
 866 and  $\sim 190$  ka (Iceland Basin, Fig. 1) may have contributed to phylogenetic nodes in  
 867 hominin evolution revealed by fossil finds and by studies of mtDNA and Y-  
 868 chromosomes (Fig. 4). According to this hypothesis, the geomagnetic field influenced  
 869 evolution of large long-lived mammals through exposure to UVR at times of low field  
 870 strength, with foci of the extinction (e.g. Australia and Europe at  $\sim 41$  ka, and North  
 871 America and Europe at  $\sim 13$  ka) depending on the specific geometry of stratospheric  
 872 ozone depletion during episodes of low field strength. Ozone holes are preferentially  
 873 located at high latitudes because of the role of stratospheric temperatures and polar

874 stratospheric clouds (PSC) in ozone depletion. UVR arriving at the Earth's surface may  
875 have had an influence on evolution due its strong mutagenic effects, its potential for  
876 promoting oxidative damage on membranes and proteins, and the role of AhR in  
877 translocating UVB stress response to the nucleus. Lower levels of UVR reaching Earth's  
878 surface at low latitudes, due to the role of polar-stratospheric clouds and stratospheric  
879 temperature in ozone depletion, may partly explain the relative diversity of modern  
880 African megafauna.

881  
882 As the chronologies of both the relative paleointensity record and the paleontological  
883 record become better resolved, we speculate that unconvincing efforts to establish a  
884 linkage between polarity reversal and extinction over the last ~50 years will, in future,  
885 be superseded by a linkage between extinction and paleointensity minima particularly  
886 for land-dwelling mammals at higher latitudes where the magnetic field strength is  
887 linked to loss of stratospheric ozone and enhanced UVR reaching Earth's surface. The  
888 importance of Australian extinction events at the time of Laschamp excursion (41 ka)  
889 and the Blake excursions (107 and 120 ka) may imply that stratospheric ozone  
890 depletion was largely in the Southern Hemisphere at these times, and in the Northern  
891 Hemisphere when the North American extinctions are manifested at ~13 ka.

892  
893 The magnetic dipole moment has decreased by ~10% since 1833 (Gauss' first direct  
894 field intensity measurement), or ~5% per century. Although the present dipole field  
895 strength may not be appreciably lower than the average during the Brunhes Chron, the  
896 current field intensity decrease, combined with asymmetry in the field detected in  
897 satellite data, has led to speculations that the geomagnetic field may reach intensity  
898 levels appropriate for a magnetic excursion or polarity reversal in the next 1000-2000  
899 years (Hulot et al., 2002; Glassmeier et al., 2009; Laj & Kissel, 2015).

900  
901 Twenty years ago, no more than 5 Quaternary magnetic excursions had been recorded  
902 (see Opdyke & Channell, 1996), whereas today ~20 Quaternary magnetic excursions  
903 have credible documentation (Laj & Channell, 2015). In the next few decades, we  
904 speculate that integration of excursion records with RPI and cosmogenic isotope  
905 records, understanding of the role of the geomagnetic field in controlling UVR flux and  
906 the role of Ahr in modulating the deleterious effects of UVR in extant and fossil  
907 mammals, documentation of mammalian extinction, and improvements in human  
908 phylogeny from mtDNA and Y-chromosomes, will clarify the role of the geomagnetic  
909 field in mammalian evolution.

910  
911  
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915  
916

917 **Figure Captions**

918

919 Fig. 1. Relative paleointensity (RPI) calibrated to virtual axial dipole moment (VADM).  
 920 (a) Overall RPI stack of 23 sedimentary records (red) with standard error ( $2\sigma$ )  
 921 compared with the  $^{10}\text{Be}$ -based VADM from Greenland ice cores (blue) filtered with cut-  
 922 offs of 1/3000 yrs (dark blue) and 1/500 yrs (dashed light blue) (Channell et al., 2018).  
 923 Also shown: Holocene VADM models of Constable et al. (2016) (green) and Pavón-  
 924 Carrasco et al. (2014) (black). (b) PISO RPI stack for 40-300 ka (Channell et al., 2009).  
 925 Yellow shading indicates RPI minima in the stack that have been associated with  
 926 directional magnetic excursions. See text for references that document the labeled  
 927 excursions.

928

929 Fig. 2. Geomagnetic field intensity (VADM) for the last 70 kyr compared with timing of  
 930 continent-wide extinction events for genera from Australia (Saltré et al., 2016), Eurasia  
 931 (Cooper et al., 2015), and North America (Faith & Surovell, 2009), including extinction  
 932 of *Arctodus simus* (Schubert, 2010) and population decline for *Bison priscus* (Shapiro et  
 933 al., 2004). Solid and dashed bars in the Eurasia record refer to fossil and DNA derived  
 934 events, respectively. Last appearance datums (LADs) for extinct North American genera  
 935 (Faith & Surovell, 2009) were recalibrated using Calib 7.1 (Stuiver et al., 2018) with  $2\sigma$   
 936 errors for entries that pass selection criteria established by Meltzer & Mead (1985).  
 937 Dung fungi (*Sporormiella* and *Podospora*) in sedimentary archives, proxies for herbivore  
 938 population, are for North America (Gill et al., 2009) and Australia (Johnson et al., 2015;  
 939 van der Kaars et al., 2017). Geomagnetic field intensity during the last 70 kyr (red)  
 940 combines VADM for the last 14 kyr (Pavón-Carrasco et al., 2014), VADM for 14-45 ka  
 941 (Channell et al., 2018), and the PISO VADM record (Channell et al., 2009) for 45-70 ka.  
 942 The light blue band represents the VADM estimate from  $^{36}\text{Cl}$ -flux in Greenland ice-cores  
 943 (Muscheler et al., 2005). Dark blue line corresponds to the ice-core  $^{10}\text{Be}$ -derived VADM  
 944 record (Channell et al., 2018). Background shades of blue are scaled to the maximum  
 945 (darker blue color) and minimum (white) VADM intensity in the  $1.35\text{-}13.27 \times 10^{22} \text{ Am}^2$   
 946 range.

947

948 Fig. 3. Geomagnetic field intensity (VADM) for the 35-50 ka interval compared to  
 949 probability density functions for the timing of Neanderthal disappearance at different  
 950 sites (blue) and for the end of the Mousterian tool industry (black) (Higham et al.,  
 951 2014), with dung-fungal proxies for large herbivore extinctions from NE Queensland  
 952 (green, Johnson et al., 2015) and southern Western Australia (open circles joined by  
 953 purple line, van der Kaars et al., 2017). The yellow band represents the GLOPIS  
 954 paleointensity record converted to VADM (Laj et al., 2004, 2014), the purple curve is the  
 955 sedimentary VADM record with  $2\sigma$  errors, and the red line is the VADM reconstructed  
 956 from  $^{10}\text{Be}$  flux in Greenland ice-cores (Channell et al., 2018). The gray shading indicates  
 957 the age range associated with Neanderthal demise, and the temporal overlap between  
 958 the two human groups (Neanderthals and Cro Magnon). The end of the Mousterian flint-  
 959 tool industry has been dated at 39.260-41.030 ka (Higham et al., 2014), and 39.528-  
 960 41.013 ka (Cooper et al., 2015), corresponding closely with low field intensity  
 961 associated with the Laschamp geomagnetic excursion (39.7-41.9 ka) (Laj et al., 2014).  
 962

963

964 Fig. 4. (a) Schematic evolutionary history of *Homo sapiens* modified after Lahr (2016).  
 965 The fossil record includes key fossil specimens from Eliye Springs, Guomde, Florisbad,  
 Herto, Omo Kibish and hominins from Jebel Irhoud dated by Richter et al., (2017).

966 Black/gray lines represent modern humans, dark blue lines the Denisovans, and the  
967 light blue lines the Neanderthals. Neanderthal extinction has been drawn according to  
968 the timing of the end of the Mousterian tool industry at ~40 ka (Higham et al., 2014).  
969 (b) Estimates of the time to the most recent common ancestor (TMRCA) and  
970 phylogenetic nodes, from Y-chromosomes (blue icons) and mitochondrial DNA  
971 (mtDNA) (pink icons), according to different authors. For Scozzari et al. (2014),  
972 haplogroup B corresponds to the split between chromosomes found only in central-  
973 west Africa and chromosomes spread over sub-Saharan Africa whereas the node at 115  
974 ka marks the separation between African specific and all remaining haplogroups. Data  
975 from Wei et al. (2013) (GENETREE-2 model) include the TMRCA of haplogroup A3 and  
976 three phylogenetic nodes considered by the authors of particular interest: DR  
977 (expansion of Y-chromosomes following the out-of-Africa migration), FR (paleolithic  
978 male lineage expansion), and R1b (Neolithic modern European chromosomes). The data  
979 from Rito et al. (2013) refer to different nodes of the typical African haplogroup L0. (c)  
980 Geomagnetic VADM for the last 300 kyr (red band) as in Figure 2, using the PISO VADM  
981 record (Channell et al., 2009) to extend the record back to 300 ka. The light blue band  
982 and the dark blue line represent the VADM estimate (from  $^{36}\text{Cl}$ - and  $^{10}\text{Be}$ -derived,  
983 respectively) from ice-cores as in Figure 2. The yellow line represents the VADM  
984 determined from the GLOPIS paleointensity stack from marine sediments (Laj et al.,  
985 2004, 2014). Magnetic excursions corresponding to paleointensity minima are labeled  
986 (see Fig. 1). Background shades of blue are scaled to the maximum (darker blue color)  
987 and minimum (white) VADM intensity in the  $1.35\text{-}14.73 \times 10^{22} \text{ Am}^2$  range.  
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1901 **GLOSSARY**

1902

1903 **6-4 photoproducts (6-4 PPs):** 6-4 pyrimidine-pyrimidone photoproducts (6-4 PPs) and  
 1904 cyclobutane pyrimidine dimers (CPDs, see below) are the common UVR products formed in  
 1905 the human skin during exposure to sunlight. Although the 6-4 PPs are not as prevalent as  
 1906 CPDs (in lesion formation), they are significantly more mutagenic.

1907 **Anhyseretic remanent magnetization (ARM):** A laboratory-induced remanent  
 1908 magnetization imparted by a strong alternating magnetic field in the presence of a weak  
 1909 direct current bias field. The bias field usually has an intensity comparable to the  
 1910 geomagnetic field. ARM provides an indication of magnetic mineral concentrations, and the  
 1911 efficiency of ARM acquisition varies with magnetic grain size.

1912 **Apoptosis:** A form of programmed cell death that occurs in multicellular organisms. It is  
 1913 a highly regulated and controlled process that confers advantages during an organism's  
 1914 lifecycle. Because apoptosis cannot be halted once initiated, it is a highly regulated process.

1915 **Artiodactyla:** Order of even-toed ungulates, hoofed animals that bear their weight on  
 1916 two toes including hippopotamuses, camels, antelopes, sheep, giraffes and deer.

1917 **Aryl hydrocarbon receptor (AhR):** a protein that regulates gene expression, cell  
 1918 physiology and organ homeostasis with a key role in skin integrity and immunity. Its  
 1919 functions are related to cell proliferation, adhesion and migration as well as to cell  
 1920 differentiation. It has a strong function in mediating the response to xenobiotic and toxins  
 1921 particularly Dioxins, and to UVR.

1922 **Brunhes Chron:** The most recent time interval of the geomagnetic polarity timescale,  
 1923 characterized by a normal polarity of the geomagnetic field, that has existed since ~773 ka.

1924 **Campanian Ignimbrite (CI):** The Campanian Ignimbrite (CI) eruption was located near  
 1925 Napoli (Italy) and was probably the largest volcanic eruption in the Mediterranean region  
 1926 during the last 200 kyr. CI ash was deposited on the Russian Plain, and throughout the  
 1927 Eastern Mediterranean and northern Africa. The event coincided with the onset of a cold  
 1928 climatic phase known as Heinrich Stadial 4 (HS4) approximately 40,000 years ago.

1929 **Chlorofluorocarbons (CFCs):** an organic compound (commonly known as Freon) that  
 1930 contains carbon, chlorine, and fluorine, produced as a volatile derivative of methane and  
 1931 ethane. Many CFCs have been commonly used as refrigerants, propellants (in aerosol  
 1932 applications), and solvents. Although the concentration of CFCs in the atmosphere is very  
 1933 small, measured in parts per trillion, they do contribute significantly to the enhancement of  
 1934 the greenhouse effect. CFCs contribute significantly to ozone depletion in the upper  
 1935 atmosphere. The manufacture of such compounds has been phased out under the Montreal  
 1936 Protocol, and their use is being replaced by other products such as hydrofluorocarbons.

1937 **Clades:** In taxonomy, a clade is defined as a group of organisms consisting of a single  
 1938 common ancestor and all descendants of that ancestor.

1939 **Clovis horizon:** The short-lived Clovis horizon comprises distinctive stone tools  
 1940 marking early human occupation of North America at ~13 ka. The Clovis culture is a  
 1941 prehistoric Paleo-Indian culture, named after characteristic stone tools found in close  
 1942 association with Late Pleistocene fauna at Blackwater near Clovis, New Mexico, in the 1920s  
 1943 and 1930s. Most of the indigenous cultures of the Americas are considered to descend from  
 1944 the Clovis people.

1945 **Coercivity:** The coercivity, also called magnetic coercivity, is a measure of the ability of  
 1946 a remanent magnetization to withstand an external demagnetizing magnetic field. For  
 1947 ferromagnetic materials (*sensu lato*), the coercivity (or coercive force) is the intensity of the  
 1948 applied magnetic field required to reduce the remanent magnetization of that material to zero  
 1949 after the magnetization of the sample has been driven to saturation in the opposite direction.

1950 Thus, coercivity measures the resistance of a ferromagnetic material (*sensu lato*) to  
 1951 demagnetization by an external field.

1952 **Cosmogenic isotopes:** Radioactive isotopes created in the upper atmosphere when  
 1953 galactic cosmic rays collide with atmospheric molecules at high speed. The production rate of  
 1954 these isotopes depends on the intensity of the cosmic radiation, which is related to the  
 1955 strength of the Earth magnetic field and solar activity. Therefore, records of cosmogenic  
 1956 isotope production are useful for understanding the relation between the Earth magnetic field,  
 1957 and variations in solar activity.

1958 **Cyclobutane pyrimidine dimers (CPDs):** One of two major photoproducts generated  
 1959 by UV irradiation. CPDs are molecular lesions formed in the human skin during exposure to  
 1960 sunlight due to UV radiation and photochemical reactions.

1961 **Cytosol:** It is the fluid inside living cells. Proteins, organelles, and other structures of the  
 1962 cells live in this water-based fluid.

1963 **Deoxyribosyl:** A univalent radical derived from deoxyribose, a monosaccharide (simple  
 1964 sugar). It is derived from the sugar as indicated by its name: deoxy sugar. Deoxyribose is the  
 1965 five-carbon sugar molecule involved to form the phosphate backbone of DNA molecules.  
 1966 DNA, or deoxyribonucleic acid, is a polymer formed of several nucleic acids. Each nucleic  
 1967 acid is composed by a molecule of deoxyribose bound to a phosphate group and either a  
 1968 purine or a pyrimidine. As part of DNA, 2-deoxyribose derivatives have a significant role in  
 1969 biology. The DNA molecule, which is the main repository of genetic information in life,  
 1970 consists of a long chain of deoxyribose-containing units called nucleotides, linked via  
 1971 phosphate groups.

1972 **DNA photoproducts:** Exposure to the ultraviolet component of sunlight causes DNA  
 1973 damage, which subsequently leads to mutations, cellular transformation, and cell death  
 1974 through the creation of different photoproducts. There are two important DNA  
 1975 photoproducts, namely cyclobutane pyrimidine dimers (CPDs) and pyrimidine-pyrimidone  
 1976 photoproducts (6-4 PPs).

1977 **Endogenous ligand:** A ligand is a protein that attaches (binds) to another protein called  
 1978 a receptor; the latter have specific sites into which the ligands fit like keys into locks. Ligands  
 1979 that are produced in the body are called endogenous.

1980 **Energetic particle precipitation (EPP):** Precipitation of a group of highly energetic  
 1981 electrons, protons, neutrons, and ions that are accelerated into the atmosphere through various  
 1982 heliophysical and geomagnetic processes. They enter the atmosphere mainly in the vicinity of  
 1983 the geomagnetic poles. When energetic particles enter the atmosphere they ionize and  
 1984 dissociate atmospheric constituents, resulting in the formation of reactive nitrogen oxides.  
 1985 EPPs have been shown to contribute up to 10% of the stratospheric NO<sub>x</sub> budget and up to  
 1986 40% of the polar stratospheric NO<sub>x</sub> budget. Once in the stratosphere, NO<sub>x</sub> produced by EPPs  
 1987 (EPP-NO<sub>x</sub>) interferes with catalytic cycles involving ozone (O<sub>3</sub>). Theoretically, changes in O<sub>3</sub>  
 1988 can also lead to changes in temperature and winds, which means that EPPs have the potential  
 1989 to impact climate.

1990 **Gaussian-resampled, inverse-weighted McInerny et al. (GRIWM):** Dating method  
 1991 for estimating the probability of extinction by using an approach that weighs observations  
 1992 inversely according to their temporal distance from the last observation of a species'  
 1993 confirmed occurrence. For dates with associated radiometric errors it is able to sample  
 1994 individual dates from an underlying fossilization probability distribution.

1995 **Geomagnetic excursions:** Excursions represent a significant, but short-lived, change of  
 1996 direction of the Earth's magnetic field, and are apparently manifested globally. Intervals of  
 1997 constant polarity are punctuated by geomagnetic excursions where magnetic directions depart  
 1998 from the usual geocentric axial dipole, and when adequately recorded, achieve the opposite  
 1999 polarity direction for a short time. Excursions have been recorded in both volcanic and

2000 sedimentary records. These directional aberrations typically have durations of a few thousand  
 2001 years, or less than 1 kyr in some cases, and are characterized by a decrease in strength of the  
 2002 main axial dipole.

2003 **Halons:** a group of hydrocarbon compounds in which some hydrogen atoms are replaced  
 2004 by bromine and fluorine atoms, and sometimes also chlorine. Halons are nonconductors of  
 2005 electricity and may be utilized in fighting fires in ignitable liquids and most solid flammable  
 2006 materials. Halons in the atmosphere are responsible for ozone depletion, and contribute to  
 2007 greenhouse warming.

2008 **Haplogroup:** group of individuals sharing a common ancestor identified by DNA  
 2009 sequences defined by shared mutations and which tend to show regional specificity. In  
 2010 human genetics, the most commonly studied haplogroups are Y-chromosome (Y-DNA)  
 2011 haplogroups and mitochondrial DNA (mtDNA) haplogroups, each of which can be used to  
 2012 describe genetic populations.

2013 **Haplotype:** A group of genes in the chromosome of an organism that are inherited  
 2014 together, because of genetic linkage from a single parent. The word "haplotype" is derived  
 2015 from the word "haploid," that describes cells with only one set of chromosomes, and from the  
 2016 word "genotype," which refers to the genetic makeup of an organism.

2017 **Heinrich Stadial (HS) 4:** A Heinrich event is a natural phenomenon occurring during a  
 2018 period known as Marine Isotops Stage 3 (about 59-27 ka BP). Large icebergs containing rock  
 2019 mass eroded by the glaciers break off from glaciers and traverse the North Atlantic, as they  
 2020 melted this material was deposited on the sea floor as ice rafted debris (IRD). One of the  
 2021 strongest HS is the number 4 originated around 40 ka BP.

2022 **Isothermal remanent magnetization (IRM):** IRM is an artificial (usually laboratory  
 2023 acquired) remanent magnetization acquired by applying a strong direct field to a  
 2024 ferromagnetic (*sensu lato*) material. It is a useful parameter used for detection of magnetic  
 2025 minerals with high coercivity such as hematite or goethite.

2026 **Late Quaternary Extinction (LQE):** Widespread extinction of large terrestrial  
 2027 mammals (megafauna) during the Late Quaternary, observed in the Americas, Europe,  
 2028 Africa, Asia and Australia. As the ages of extinction have become better constrained, the  
 2029 timing of these extinctions has become focused in the vicinity of 40 ka and 13 ka.

2030 **Ligand-binding:** Intermolecular interactions occurring between or among proteins,  
 2031 nucleic acids, or small molecules. The interaction of ligands with their binding sites are often  
 2032 characterized in terms of a binding affinity. In general, high-affinity ligand binding occurs  
 2033 from greater intermolecular force between the ligand and its receptor, whereas low-affinity  
 2034 ligand binding involves less intermolecular force between the ligand and its receptor.

2035 **Luminescence age dating:** refers to a group of dating methods based on the determining  
 2036 of how long ago mineral grains were last exposed to sunlight. The techniques include  
 2037 optically stimulated luminescence (OSL), infrared stimulated luminescence (IRSL), and  
 2038 thermoluminescence dating (TL). It is useful to geologists and archaeologists who want to  
 2039 know the age of a particular event. Sediments and soils contain traces of radioactive isotopes  
 2040 such as potassium, uranium, thorium, and rubidium. The ionizing radiation emitted during the  
 2041 decay of these elements is absorbed and trapped by quartz and potassium feldspar grains  
 2042 occurring in the sediment. Stimulating these mineral grains using either light (blue or green  
 2043 for OSL; infrared for IRSL) or heat (for TL) produces a luminescence signal emitted as  
 2044 stored unstable electron energy, the intensity of which varies depending on the amount of  
 2045 radiation absorbed during burial, and specific properties of the mineral.

2046 **Mediterranean sapropel S7:** The sediments of the Mediterranean (central-eastern) are  
 2047 characterized by the presence of organic-rich (>2% organic carbon) layers called sapropels.  
 2048 They were deposited during periods of reduced oxygenation related to increasing riverine  
 2049 input occurring in correspondence with peaks in solar insolation. Each peak has been

2050 numbered starting from the youngest (S1; 8-10 ka) and the sapropel S7 occurred around 190  
2051 ka.

2052 **Melanocyte homeostasis:** The maintenance of systems within a cell as conditions  
2053 change is called homeostatic regulation, and melanocyte (a mature melanin-forming skin  
2054 cell) homeostasis is a paradigm for understanding the formation of melanoma.

2055 **Mitochondria:** Mitochondria are part of a eukaryote cell found in the cytoplasm. They  
2056 oxidise glucose to provide energy for the cell through the creation of a molecule, called  
2057 adenosine triphosphate that cells use as an energy source.

2058 **Mitochondrial DNA (mtDNA):** Mitochondria are tiny organelles that live in the  
2059 cytoplasm of cells, called mitochondria. Each cell contains thousands of mitochondria with  
2060 its own small circle of DNA, a reminder of their distant bacterial ancestry. MtDNA is not  
2061 located in the cell nucleus and in most species, including humans, it is inherited solely from  
2062 the mother and is not subject to genetic recombination during meiosis. Therefore it remains  
2063 unchanged from generation to generation. Human mtDNA was the first significant part of the  
2064 human genome to be sequenced. Since animal mtDNA evolves faster than nuclear genetic  
2065 markers, it represents a pillar of phylogenetics and evolutionary biology. It also makes  
2066 possible an examination of the affinity of populations, and so has become important in  
2067 anthropology and biogeography.

2068 **Mitochondrial Eve.** Analyses of the mitochondrial DNA of living humans have shown  
2069 that we descend from a common female ancestor that has been dated to about 200 ka. This  
2070 woman has been named "Eve" or "Mitochondrial Eve" after a journalist's confusing reference  
2071 to the unrelated biblical story of the first woman being created by God from Adam's rib.  
2072 ("The man called his wife Eve ["life"], because she was the mother of all who live." Genesis  
2073 3:20)

2074 **Mitohormesis:** Mitohormesis, known also as mitochondrial hormesis, is a particular  
2075 form of hormesis that is a non-linear response to potentially harmful substances.  
2076 Mitochondrial function often results in excessive production of reactive oxygen species  
2077 (ROS) that are responsible for many chronic diseases. However, moderate levels of  
2078 mitochondrial ROS, can protect against chronic disease by stimulating mitochondrial  
2079 capacity and endogenous antioxidant defence. This phenomenon is called mitohormesis.

2080 **Mousterian tool industry:** The predominant industry of the Middle Paleolithic is  
2081 termed Mousterian, named for its type-site Le Moustier, in Dordogne, France. Mousterian  
2082 industry is the tool culture traditionally associated with Neanderthal man in Europe, western  
2083 Asia, and northern Africa. Mousterian tools disappeared abruptly from Europe with the  
2084 passing of Neanderthal man.

2085 **Natural remanent magnetization (NRM):** The magnetic remanence carried by rocks or  
2086 sediments before the laboratory demagnetization or magnetization treatments. The NRM of a  
2087 rock or sediment is usually represented by multiple magnetization components acquired  
2088 through its geologic history. Thermal or alternating field laboratory treatments are usually  
2089 necessary to identify the primary remanence acquired at the time of the rock formation or  
2090 sediment deposition.

2091 **Pedigree-based substitution rate:** A method to estimate mutation rate by comparing the  
2092 mtDNA sequences of a sample of parent/offspring pairs or analyzing mtDNA sequences of  
2093 individuals from a deep-rooted (well established) genealogy. The number of new mutations  
2094 in the sample is counted and divided by the total number of parent-to-child DNA  
2095 transmission events to calculate a mutation rate.

2096 **Perissodactyla:** Order of odd-toed ungulates, hoofed animals that bear most of their  
2097 weight on a single toe including horses, zebras, rhinoceroses and tapirs.

2098 **Peto's paradox:** Peto's paradox, named after Richard Peto, is the observation that at  
2099 species level, the incidence of cancer does not appear to correlate with the number of cells in

2100 an organism (e.g., body weight). Although cell mutations are usually deleterious, the  
 2101 incidence of cancer is not apparently related to the body-size of the individual mammals.

2102 **Phenotypical.** The observable physical or biochemical characteristics of an organism  
 2103 determined by genetic and environmental influences.

2104 **Phylogeny:** The evolutionary history and relationships among people or groups of  
 2105 organisms (e.g. species or populations). These relationships can be discovered through  
 2106 phylogenetic methods that evaluate observed heritable traits, such as DNA sequences or  
 2107 morphology, using an evolutionary model for these traits. The results of these analyses  
 2108 represent the phylogeny (also referred to as the phylogenetic tree).

2109 **Polar-stratospheric clouds:** Polar-stratospheric clouds are typically formed in the  
 2110 stratosphere (15-25 km altitude) at high latitudes especially in winter because they require  
 2111 very low temperatures. They are implicated in the formation of ozone holes because they  
 2112 convert benign forms of chlorine into reactive, ozone-destroying forms, and they remove  
 2113 gaseous nitric acid that would otherwise moderate the destructive impact of chlorine.

2114 **Quaternary:** The most recent of the three periods of the Cenozoic Era in the geologic  
 2115 time scale. It spans the last 2.588 Myrs and is divided into two epochs: the Pleistocene (2.588  
 2116 million years ago to 11.7 thousand years ago) and the Holocene (11.7 thousand years ago to  
 2117 present).

2118 **Reactive oxygen species (ROS):** Oxygen-derived molecules that act as powerful  
 2119 oxidants and may form via a large number of physiologic and non-physiologic processes as a  
 2120 result of natural consequences of aerobic metabolism. Although these molecules play a role  
 2121 in the oxygen-dependent defence mechanism against bacteria, they may also be highly  
 2122 damaging, as they can attack biologic macromolecules, including lipids, proteins, and DNA,  
 2123 and lead to significant tissue damage. In normal conditions, ROS can act as immune system  
 2124 modulation and can activate various signal transduction pathways. ROS levels can increase  
 2125 dramatically during times of environmental stress (e.g., UVR or heat exposure).

2126 **Relative paleointensity (RPI):** The intensity of the Earth's axial dipole field is globally  
 2127 coherent, and can be reconstructed in the past by studying suitable rock and sediment samples  
 2128 or archaeological materials. The RPI proxy in sediments usually comprises the NRM  
 2129 intensity normalized by the intensity of a laboratory-acquired magnetization designed to  
 2130 activate the same population of magnetic grains that carry the NRM, thus compensating for  
 2131 variations in concentration of NRM-carrying grains through the sedimentary section. The  
 2132 laboratory-applied normalizers are typically anhysteretic remanent magnetization (ARM)  
 2133 and/or isothermal remanent magnetization (IRM).

2134 **Signor-Lipps effect:** The first and last occurrence of a taxon will never be accurately  
 2135 recorded in the fossil record because of the inherent incompleteness of the record.

2136 **Sclerophyll vegetation:** Type of vegetation characterized by hard, leathery, evergreen  
 2137 foliage well adapted to prevent moisture loss. Broad-leaved sclerophyll vegetation, including  
 2138 species such as holly (*Ilex*), is known as Mediterranean vegetation because it develops in  
 2139 regions with a Mediterranean climate - hot, dry summers and mild, wet winters. Pines are  
 2140 examples of narrow-leaved sclerophyll vegetation.

2141 **Solar proton events (SPE):** A SPE or "proton storm", occurs when particles (mostly  
 2142 protons) emitted by the Sun are accelerated either close to the Sun during a flare or in  
 2143 interplanetary space by coronal mass ejection (CME). Solar protons normally have  
 2144 insufficient energy to penetrate the Earth's protective magnetosphere. However, during  
 2145 unusually strong flares, protons can be produced with sufficient energies to penetrate the  
 2146 Earth's magnetosphere and ionosphere around the poles. They can pass through the Earth's  
 2147 magnetic field and cause ionization in the ionosphere. The effect is analogous to an auroral  
 2148 event, except that protons instead of electrons are involved. Energetic protons represent a  
 2149 significant radiation hazard to spacecraft and astronauts.

2150 **The time to the most recent common ancestor (TMRCAs):** Time to the most  
2151 recent individual from which an entire organism is directly descended. The term is used in  
2152 reference to the ancestry of groups of genes (haplotypes) or species.

2153 **TP53 (p53) gene:** Tumour protein 53 is a protein-coding gene that acts as a tumour  
2154 suppressor for many tumour types, and induces growth arrest or apoptosis depending on  
2155 physiological circumstances and cell type. The TP53 gene is responsible for instructions to  
2156 create a protein called tumour protein p53. This protein acts as a tumour suppressor, which  
2157 means that it regulates cell division by keeping cells from growing and dividing  
2158 (proliferating) too fast or in an uncontrolled way. The p53 protein is found in the nucleus of  
2159 cells throughout the body, where it attaches (binds) directly to DNA. If the DNA in a cell is  
2160 damaged by agents such as toxic chemicals, radiation or UVR, this protein plays a critical  
2161 role as to whether the DNA can be repaired or the damaged cell will self-destruct (undergo  
2162 apoptosis). If it is possible to repair the DNA, p53 activates other genes to fix the damage. On  
2163 the other hand, this protein prevents the cell from dividing and signals it to undergo  
2164 apoptosis. By stopping cells with mutated or damaged DNA from dividing, p53 helps prevent  
2165 tumour development.

2166 **Tryptophan:** Tryptophan is a  $\alpha$ -amino acid that is used in the biosynthesis of proteins.  
2167 Many animals (including humans) cannot synthesize tryptophan: they obtain it through their  
2168 diet. Tryptophan is among the less common amino acids found in proteins, but it has critical  
2169 structural and functional roles.

2170 **Virtual axial dipole moment (VADM):** Intensity of an imaginary axial (along the  
2171 Earth's rotation axis) centric (located in the centre of the Earth) dipole that would produce the  
2172 estimated magnetic field reconstructed from paleomagnetic data. The reconstruction is virtual  
2173 because the VADM is an approximation of a geocentric axial dipole moment.

2174 **Xenobiotics:** A xenobiotic is a chemical substance within an organism that is foreign to  
2175 the biological system, and not naturally produced or expected to be present within the  
2176 organism. It can also cover substances that are present in much higher concentrations than  
2177 usual. The term is often used in the context of pollutants, and their effect on biota.

2178 **Y-chromosome:** one of the two sex chromosomes in humans (the other is the X  
2179 chromosome) and most mammals. The Y chromosome occurs in males, who have one X and  
2180 one Y chromosome, while females have two X chromosomes. The Y chromosome contains  
2181 genes that provide instructions for making proteins. Because only males have the Y  
2182 chromosome, the genes in this chromosome are involved in male sex determination and  
2183 development. The Y chromosome represents almost 2 % of the total DNA in male cells.

2184  
2185

Figure 1.